

Non-Hermitian Quantum Systems and Time-Optimal Quantum Evolution^{*}

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Abstract. Recently, Bender et al. have considered the quantum brachistochrone problem for the non-Hermitian \mathcal{PT} -symmetric quantum system and have shown that the optimal time evolution required to transform a given initial state $|\psi_i\rangle$ into a specific final state $|\psi_f\rangle$ can be made arbitrarily small. Additionally, it has been shown that finding the shortest possible time requires only the solution of the two-dimensional problem for the quantum system governed by the effective Hamiltonian acting in the subspace spanned by $|\psi_i\rangle$ and $|\psi_f\rangle$. In this paper, we study a similar problem for the generic non-Hermitian Hamiltonian, focusing our attention on the geometric aspects of the problem.

Key words: non-Hermitian quantum systems; quantum brachistochrone problem

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1 Introduction

In view of recent results on optimal quantum evolution and its possible relation to quantum computation and quantum information processing, there has been increasing interest in the quantum brachistochrone problem. The problem consists of finding the optimal time evolution τ to evolve a given initial state $|\psi_i\rangle$ into a certain final state $|\psi_f\rangle$ under a given set of constraints [1, 2, 3, 4, 5]. It is known that for the Hermitian Hamiltonian, τ has a nonzero lower bound. However, Bender et al. [2] have recently shown that for non-Hermitian \mathcal{PT} -symmetric quantum systems, the answer is quite different and that the time evolution τ can be made arbitrarily small, despite the fact that the eigenvalue constraint is held fixed and identical to that for the corresponding Hermitian system. The mechanism described in [2] resembles the wormhole effect in general relativity and has generated discussion in the literature [6, 7, 8, 9, 10, 11, 12, 13, 14].

Non-Hermitian Hamiltonians emerge in physics in different ways. For instance, the application for description of dissipative systems is well known from the classical works by Weisskopf and Wigner on metastable states [15, 16, 17]. It was demonstrated that the evolution of the quantum system, being initially in the metastable state $\psi(0)$, can be described by the effective non-Hermitian Hamiltonian H_{eff} as follows: $\psi(0) \rightarrow \psi(t) = e^{-iH_{\text{eff}}t}\psi(0) + \text{decay products}$. Recently, it has been shown how the non-Hermitian Hamiltonian appears in the framework of the quantum jump approach to open systems [18, 19]. Other examples include complex refractive indices in optics, complex potentials describing the scattering of electrons, atom diffraction by light, line widths of unstable lasers, etc. For the following, it is essential that non-Hermitian physics differs drastically from the conventional physics in the presence of the so-called *except-*

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tional points, where the eigenvalues and eigenvectors coalesce, even if the non-Hermiticity is regarded as a perturbation [20].

In the Hermitian quantum mechanics, the optimal time evolution problem implies finding the transformation $|\psi_i\rangle \rightarrow e^{-iHt}|\psi_i\rangle$ that provides the shortest time $t = \tau$ under a given set of constraints [1]. The generic non-Hermitian quantum brachistochrone problem poses the same question, with the exception that now the evolution of the system is described by the non-Hermitian Hamiltonian, which is not necessarily \mathcal{PT} -symmetric [2, 6, 12].

Recently, this problem has been studied by Assis and Fring for a dissipative quantum system governed by a symmetric non-Hermitian Hamiltonian [6]. It has been shown that the passage time required to transform a given initial state to the orthogonal final state can be arbitrarily small. The obtained effect, being similar to the one discovered by Bender et al. [2], is related to the non-Hermitian nature of the system rather than to its \mathcal{PT} -symmetry.

In this paper, we address the non-Hermitian quantum brachistochrone problem, considering the generic non-Hermitian Hamiltonian, and focus our attention on the geometric aspects of the problem. In Section 2 we first briefly summarize the main properties of non-Hermitian quantum systems. We discuss the non-Hermitian quantum brachistochrone problem and the associated two-dimensional effective Hamiltonian, acting in the two-dimensional space spanned by the initial and final states. In Section 3 we discuss the quantum brachistochrone in the vicinity of the exceptional point. In Section 4 we consider the non-Hermitian quantum brachistochrone problem in the context of the Fubini–Study metric on the complex Bloch sphere. In Section 5 the results and open problems are discussed.

2 Non-Hermitian quantum brachistochrone problem

In the quantum brachistochrone problem, finding the shortest possible time requires only the solution of the two-dimensional problem, namely, finding the optimal time evolution for the quantum system governed by the effective Hamiltonian acting in the subspace spanned by $|\psi_i\rangle$ and $|\psi_f\rangle$ [2, 3, 4].

Before proceeding further, we outline some background information on non-Hermitian quantum systems. Let an adjoint pair $\{|\psi(t)\rangle, |\tilde{\psi}(t)\rangle\}$ be a solution to the Schrödinger equation and its adjoint equation

$$i\frac{\partial}{\partial t}|\psi(t)\rangle = H|\psi(t)\rangle, \quad i\frac{\partial}{\partial t}|\tilde{\psi}(t)\rangle = H^\dagger|\tilde{\psi}(t)\rangle,$$

which can be recast in the form

$$i\frac{\partial}{\partial t}|\psi(t)\rangle = H|\psi(t)\rangle, \quad -i\frac{\partial}{\partial t}\langle\tilde{\psi}(t)| = \langle\tilde{\psi}(t)|H. \quad (2.1)$$

For λ_k being the eigenvalues of H , we denote by $|\psi_k\rangle$ and $\langle\tilde{\psi}_k|$ the corresponding right and left eigenvectors: $H|\psi_k\rangle = \lambda_k|\psi_k\rangle$, $\langle\tilde{\psi}_k|H = \lambda_k\langle\tilde{\psi}_k|$. The systems of both left and right eigenvectors form bi-orthogonal basis [21]

$$\sum_k \frac{|\psi_k\rangle\langle\tilde{\psi}_k|}{\langle\tilde{\psi}_k|\psi_k\rangle} = 1, \quad \langle\tilde{\psi}_k|\psi_{k'}\rangle = 0, \quad k \neq k'.$$

Let the set $\{|\psi_i\rangle, |\psi_0\rangle, \langle\tilde{\psi}_i|, \langle\tilde{\psi}_0|\}$ form the bi-orthonormal basis of the two-dimensional subspace spanned by the initial state $|\psi_i\rangle$ and the final state $|\psi_f\rangle$:

$$\langle\tilde{\psi}_0|\psi_0\rangle = \langle\tilde{\psi}_i|\psi_i\rangle = 1, \quad \langle\tilde{\psi}_0|\psi_i\rangle = \langle\tilde{\psi}_i|\psi_0\rangle = 0.$$

Using this basis, we can write the final state $|\psi_f\rangle$ as

$$|\psi_f\rangle = \cos \frac{\alpha}{2} |\psi_i\rangle + e^{i\beta} \sin \frac{\alpha}{2} |\psi_0\rangle, \quad (2.2)$$

where α, β are complex angles, and similarly,

$$\langle \tilde{\psi}_f | = \cos \frac{\tilde{\alpha}}{2} \langle \tilde{\psi}_i | + e^{-i\tilde{\beta}} \sin \frac{\tilde{\alpha}}{2} \langle \tilde{\psi}_0 |.$$

Let $|\psi(t)\rangle = u_1(t)|\psi_i\rangle + u_2(t)|\psi_0\rangle$ and $\langle \tilde{\psi}(t) | = \tilde{u}_1 \langle \tilde{\psi}_i | + \tilde{u}_2 \langle \tilde{\psi}_0 |$ be solutions to the Schrödinger equation and its adjoint equation (2.1), respectively. Then one can see that the vectors $|u(t)\rangle = \begin{pmatrix} u_1(t) \\ u_2(t) \end{pmatrix}$ and $\langle \tilde{u}(t) | = (\tilde{u}_1(t), \tilde{u}_2(t))$ satisfy

$$i \frac{\partial}{\partial t} |u(t)\rangle = H_{\text{eff}} |u(t)\rangle, \quad -i \frac{\partial}{\partial t} \langle \tilde{u}(t) | = \langle \tilde{u}(t) | H_{\text{eff}}, \quad (2.3)$$

where the effective two-dimensional Hamiltonian H_{eff} reads:

$$H_{\text{eff}} = \frac{1}{2} \begin{pmatrix} \lambda_0 + Z & X - iY \\ X + iY & \lambda_0 - Z \end{pmatrix}.$$

Here, we set

$$\begin{aligned} \lambda_0 &= \langle \tilde{\psi}_i | H | \psi_i \rangle + \langle \tilde{\psi}_0 | H | \psi_0 \rangle, & X &= \langle \tilde{\psi}_i | H | \psi_0 \rangle + \langle \tilde{\psi}_0 | H | \psi_i \rangle, \\ Y &= i(\langle \tilde{\psi}_i | H | \psi_0 \rangle - \langle \tilde{\psi}_0 | H | \psi_i \rangle), & Z &= \langle \tilde{\psi}_i | H | \psi_i \rangle - \langle \tilde{\psi}_0 | H | \psi_0 \rangle. \end{aligned}$$

Further, it is convenient to express the effective Hamiltonian H_{eff} in terms of the Pauli matrices:

$$H_{\text{eff}} = \frac{\lambda_0}{2} \mathbf{1} + \frac{1}{2} \mathbf{\Omega} \cdot \boldsymbol{\sigma}, \quad (2.4)$$

where $\mathbf{1}$ denotes the identity operator and $\mathbf{\Omega} = (X, Y, Z)$.

It is then easy to show that the complex Bloch vector defined as $\mathbf{n}(t) = \langle \tilde{u}(t) | \boldsymbol{\sigma} | u(t) \rangle$, where $\boldsymbol{\sigma}$ are the Pauli matrices, satisfies the complex Bloch equation

$$\frac{d\mathbf{n}(t)}{dt} = \mathbf{\Omega} \times \mathbf{n}(t),$$

which is equivalent to the Schrödinger equation (2.3) (see, e.g., [22, 23]). In the explicit form, the components of the Bloch vector are given by

$$n_1 = u_1 \tilde{u}_2 + u_2 \tilde{u}_1, \quad n_2 = i(u_1 \tilde{u}_2 - u_2 \tilde{u}_1), \quad n_3 = u_1 \tilde{u}_1 - u_2 \tilde{u}_2. \quad (2.5)$$

The vector $\mathbf{n}(t) = (n_1(t), n_2(t), n_3(t))$, being a complex unit vector, traces out a trajectory on the complex 2-dimensional sphere S_c^2 , and the latter can be considered as the quantum phase space for the non-Hermitian two-level quantum system. From equation (2.5), we find that $\mathbf{n}_i = (0, 0, 1)$ corresponds to the initial state $|\psi_i\rangle$.

Analysis of the eigenvalue problem $H_{\text{eff}} |u_{\pm}\rangle = \lambda |u_{\pm}\rangle$, $\langle \tilde{u}_{\pm} | H_{\text{eff}} = \lambda_{\pm} \langle \tilde{u}_{\pm} |$, yields $\lambda_{\pm} = (\lambda_0 \pm R)/2$, where $R = \sqrt{X^2 + Y^2 + Z^2}$. The right and left eigenvectors are found to be

$$\begin{aligned} |u_+\rangle &= \begin{pmatrix} \cos \frac{\theta}{2} \\ e^{i\varphi} \sin \frac{\theta}{2} \end{pmatrix}, & \langle \tilde{u}_+ | &= \left(\cos \frac{\theta}{2}, e^{-i\varphi} \sin \frac{\theta}{2} \right), \\ |u_-\rangle &= \begin{pmatrix} -e^{-i\varphi} \sin \frac{\theta}{2} \\ \cos \frac{\theta}{2} \end{pmatrix}, & \langle \tilde{u}_- | &= \left(-e^{i\varphi} \sin \frac{\theta}{2}, \cos \frac{\theta}{2} \right), \end{aligned}$$

where

$$\cos \frac{\theta}{2} = \sqrt{\frac{R+Z}{2R}}, \quad \sin \frac{\theta}{2} = \sqrt{\frac{R-Z}{2R}}, \quad (2.6)$$

$$e^{i\varphi} = \frac{X+iY}{\sqrt{R^2-Z^2}}, \quad e^{-i\varphi} = \frac{X-iY}{\sqrt{R^2-Z^2}}, \quad (2.7)$$

and

$$X = R \sin \theta \cos \varphi, \quad Y = R \sin \theta \sin \varphi, \quad Z = R \cos \theta, \quad (2.8)$$

(θ, φ) being the complex spherical coordinates.

The coalescence of eigenvalues λ_+ and λ_- occurs when $X^2 + Y^2 + Z^2 = 0$. There are two cases. The first one, defined by $\theta = 0$, $\varphi = 0$, yields two linearly independent eigenvectors. The related degeneracy is known as the diabolic point, and we obtain

$$|u_+\rangle = \begin{pmatrix} 1 \\ 0 \end{pmatrix}, \quad \langle \tilde{u}_+| = (1, 0), \quad |u_-\rangle = \begin{pmatrix} 0 \\ 1 \end{pmatrix}, \quad \langle \tilde{u}_-| = (0, 1).$$

The second case is characterized by the coalescence of eigenvalues and the merging of the eigenvectors. The degeneracy point is known as the exceptional point, and we obtain: $|u_+\rangle = e^{i\kappa}|u_-\rangle$ and $\langle \tilde{u}_+| = e^{-i\kappa}\langle \tilde{u}_-|$, where $\kappa \in \mathbb{C}$ is a complex phase.

Let us assume that the exceptional point is given by $\mathbf{R}_0 = (X_0, Y_0, Z_0)$. Then, if $Z_0 \neq 0$, using equations (2.6)–(2.8), we obtain

$$\tan \frac{\theta_0}{2} = \pm i, \quad e^{2i\varphi_0} = \frac{X_0 + iY_0}{X_0 - iY_0};$$

thus, at the exceptional point $\Im\theta \rightarrow \pm\infty$. If $Z_0 = 0$, we obtain $X_0 = \pm iY_0$. This implies that $\theta_0 = \pi/2$, and $\Im\varphi \rightarrow \pm\infty$ at the exceptional point.

Taking the Hamiltonian of equation (2.4), we find the solution of the Schrödinger equation (2.3), satisfying $|\psi(0)\rangle = |\psi_i\rangle$, as

$$|\psi(t)\rangle = C_1(t)e^{-i\lambda_0 t/2}|\psi_i\rangle + C_2(t)e^{-i\lambda_0 t/2}|\psi_0\rangle, \quad (2.9)$$

where

$$C_1(t) = \cos \frac{\Omega t}{2} - i \cos \theta \sin \frac{\Omega t}{2}, \quad C_2(t) = -ie^{i\varphi} \sin \theta \sin \frac{\Omega t}{2}, \quad (2.10)$$

and we denote $\Omega = R = \sqrt{X^2 + Y^2 + Z^2}$.

The solution of the adjoint Schrödinger equation with the wave function $\langle \tilde{\psi}(t)|$ written as

$$\langle \tilde{\psi}(t)| = \tilde{C}_1(t)e^{i\lambda_0 t/2}\langle \tilde{\psi}_i| + \tilde{C}_2(t)e^{i\lambda_0 t/2}\langle \tilde{\psi}_0| \quad (2.11)$$

is given by

$$\tilde{C}_1(t) = \cos \frac{\Omega t}{2} + i \cos \theta \sin \frac{\Omega t}{2}, \quad \tilde{C}_2(t) = ie^{-i\varphi} \sin \theta \sin \frac{\Omega t}{2}. \quad (2.12)$$

Applying equation (2.2), we can write $|\psi(t)\rangle$ as

$$|\psi(t)\rangle = \left(C_1(t) - e^{-i\beta} \cot \frac{\alpha}{2} C_2(t) \right) e^{-i\lambda_0 t/2} |\psi_i\rangle + \frac{e^{-i\beta}}{\sin \frac{\alpha}{2}} C_2(t) e^{-i\lambda_0 t/2} |\psi_f\rangle. \quad (2.13)$$

Hence, the initial state $|\psi_i\rangle$ evolves into the final state $|\psi_f\rangle$ in the time $t = \tau$ when

$$C_1(\tau) - C_2(\tau)e^{-i\beta} \cot \frac{\alpha}{2} = 0, \quad C_2(\tau) = e^{i\beta} \sin \frac{\alpha}{2}. \quad (2.14)$$

To find the time evolution τ , we first solve the equations (2.14) for $e^{i(\varphi-\beta)}$. The computation yields

$$e^{i(\varphi-\beta)} = \frac{-\cos \theta \cos \frac{\alpha}{2} \pm \sqrt{\cos^2 \theta - \sin^2 \frac{\alpha}{2}}}{\sin \theta \sin \frac{\alpha}{2}}, \quad (2.15)$$

$$e^{-i(\varphi-\beta)} = \frac{-\cos \theta \cos \frac{\alpha}{2} \mp \sqrt{\cos^2 \theta - \sin^2 \frac{\alpha}{2}}}{\sin \theta \sin \frac{\alpha}{2}}. \quad (2.16)$$

It should be noted that the solution related to the lower sign can be obtained from the solution corresponding to the upper sign by changing the parameter α as follows: $\alpha \rightarrow \alpha + 2\pi$. This implies the change of the total sign in the final state of the quantum-mechanical system: $|\psi_f\rangle \rightarrow -|\psi_f\rangle$. Thus, without loss of generality, one can consider only one sign (upper or lower) in equations (2.15), (2.16). Further, for definiteness, we will choose the upper sign. Then, substituting the result into (2.14), we get

$$\tan \frac{\Omega\tau}{2} = \frac{i \sin^2 \frac{\alpha}{2}}{\cos \frac{\alpha}{2} \sqrt{\cos^2 \theta - \sin^2 \frac{\alpha}{2}} - \cos \theta}. \quad (2.17)$$

From here, the time evolution τ is found to be

$$\tau = \left| \frac{2}{\Omega} \arctan \left(\frac{i \sin^2 \frac{\alpha}{2}}{\cos \frac{\alpha}{2} \sqrt{\cos^2 \theta - \sin^2 \frac{\alpha}{2}} - \cos \theta} \right) \right|. \quad (2.18)$$

In addition, since τ should be a real positive function, the following restriction must be imposed:

$$\arg \Omega = \arg \arctan \left(\frac{i \sin^2 \frac{\alpha}{2}}{\cos \frac{\alpha}{2} \sqrt{\cos^2 \theta - \sin^2 \frac{\alpha}{2}} - \cos \theta} \right). \quad (2.19)$$

Now, the generic problem is to select a final vector $|\psi_f\rangle$, choosing the parameters α and β . Next, we must find the conditions that should be imposed on the parameters (θ, φ) to yield the smallest time τ required to evolve the state $|\psi_i\rangle$ into the state $|\psi_f\rangle$ under a given set of constraints with the Hamiltonian

$$H_{\text{eff}} = \frac{1}{2} \begin{pmatrix} \lambda_0 & 0 \\ 0 & \lambda_0 \end{pmatrix} + \frac{\Omega}{2} \begin{pmatrix} \cos \theta & e^{-i\varphi} \sin \theta \\ e^{i\varphi} \sin \theta & -\cos \theta \end{pmatrix}. \quad (2.20)$$

In what follows, we assume the eigenvalue constraint to be imposed as $|\lambda_+ - \lambda_-| = |\Omega|$ ¹. This implies that the argument of Ω is determined by equation (2.19).

Further study of the critical points shows that there is no solution with a finite value of $|\theta|$ yielding the minimum of the time evolution [24]. Moreover, as follows from equation (2.18), in the limit $|\theta| \rightarrow \infty$ ($|\Im\theta| \rightarrow \infty$), the time evolution behaves as $\tau \approx |2/\Omega \cos \theta| \rightarrow 0$. Thus, for a quantum-mechanical system governed by a non-Hermitian Hamiltonian, the time evolution τ

¹In contrast to the Hermitian quantum systems, here, there are more choices for imposing the eigenvalue constraint. For instance, instead of $|\lambda_+ - \lambda_-| = |\Omega|$ being held fixed, one can require $\lambda_+ - \lambda_- = \Omega = \text{const}$ [24].

can be taken to be arbitrarily small. In addition, since for any finite value of $|\theta|$, the minimum of τ does not exist, the non-Hermitian Hamiltonian cannot be optimized.

However, for a quantum-mechanical system governed by a Hermitian Hamiltonian, the latter can be optimized. Indeed, in the Hermitian case, $\Im\varphi = \Im\theta = 0$; hence, the saddle point $\theta = \pi/2$ becomes the point of a local minimum (see Fig. 1). It then follows from (2.20) that

$$H_{\text{eff}} = \frac{1}{2} \begin{pmatrix} \lambda_0 & 0 \\ 0 & \lambda_0 \end{pmatrix} + \frac{\Omega}{2} \begin{pmatrix} 0 & e^{-i\varphi} \\ e^{i\varphi} & 0 \end{pmatrix},$$

and from equations (2.13)–(2.16), we obtain

$$|\psi(t)\rangle = e^{-i\lambda_0 t/2} \left(\cos \frac{\Omega t}{2} - \cot \frac{\alpha}{2} \sin \frac{\Omega t}{2} \right) |\psi_i\rangle + \frac{e^{-i\lambda_0 t/2}}{\sin \frac{\alpha}{2}} \sin \frac{\Omega t}{2} |\psi_f\rangle.$$

This fully agrees with the results obtained by Carlini et al. and Brody and Hook [1, 4].

We now consider some illustrative examples. Of a special interest is the case when the complex Bloch vector \mathbf{n} , entering in the non-Hermitian Hamiltonian (2.4), is orthogonal to the plane spanned by \mathbf{n}_i and \mathbf{n}_f . Without loss of generality, one can choose $\mathbf{n} = \mathbf{n}_i \times \mathbf{n}_f / \sin \alpha$ and set $\mathbf{n}_i = (0, 0, 1)$. This yields $\theta = \pi/2$ and from equation (2.15), we obtain $e^{i\varphi} = ie^{i\beta}$. Applying these relations, we get

$$H_{ef} = \frac{1}{2} \begin{pmatrix} \lambda_0 & 0 \\ 0 & \lambda_0 \end{pmatrix} + \frac{\Omega}{2} \begin{pmatrix} 0 & -ie^{-i\beta} \\ ie^{i\beta} & 0 \end{pmatrix}. \quad (2.21)$$

Then, using equation (2.17), we obtain the optimal time evolution as $\tau = |\alpha/\Omega|$, and from equation (2.19), we get $\Omega = |\Omega|e^{i\arg \alpha}$. This example has a simple geometric interpretation. As can be easily shown, the final state \mathbf{n}_f is obtained from the initial state \mathbf{n}_i by rotating the complex Bloch sphere S_c^2 through the complex angle α around the axes defined by the vector \mathbf{n} . It should be noted that for $\Im\alpha = \Im\beta = 0$, we have $\Im\Omega = 0$, and the Hamiltonian (2.21) coincides with the optimal Hamiltonian, yielding the shortest time evolution τ_m for the unitary evolution. As we will show in the following sections, τ_m gives the upper bound of the time evolution for the generic non-Hermitian Hamiltonian.

The other interesting example is when the initial and final states are orthogonal to each other ($\alpha = \pi$). The smallest time τ_p required to evolve from a given initial state $|\psi_i\rangle$ to the orthogonal final state $|\psi_f\rangle$ is called the *passage time* [3, 4]. Inserting $\alpha = \pi$ into equation 2.18, we obtain

$$\tau_p = \left| \frac{i}{\Omega} \ln \frac{Z - \Omega}{Z + \Omega} \right|.$$

Let us consider some limiting cases, starting with the saddle point $Z = 0$ (see Fig. 1). Computation yields $\tau_p = \pi/|\Omega|$. This can be interpreted as the Hermitian limit of the non-Hermitian system. As can be seen in Fig. 1, the evolution of the non-Hermitian quantum system is faster than that of the corresponding Hermitian system, satisfying the same eigenvalue constraint. Moreover, the passage time $\tau_p \rightarrow 0$ while $|Z| \rightarrow \infty$.

Next, we find that $\tau_p \rightarrow \infty$ at the points $Z = \pm\Omega$ (see Fig. 1). To understand this result, we note that $Z = \Omega$ implies $\theta = 0$ and that $Z = -\Omega$ yields $\theta = \pi$. For both cases, $\theta = 0$ and $\theta = \pi$, the wave function (2.13) becomes

$$|\psi(t)\rangle = e^{-i(\lambda_0 \pm \Omega)t/2} |\psi_i\rangle.$$

Thus, only phase and amplitude of the wave function are changed; otherwise, the system in the same initial state remains intact. This explains the divergence of the passage time at the points $Z = \pm\Omega$. This is easy to understand by looking at the example of the half-spin particle in the uniform magnetic field, when the spin is equinarrayed with the magnetic field.

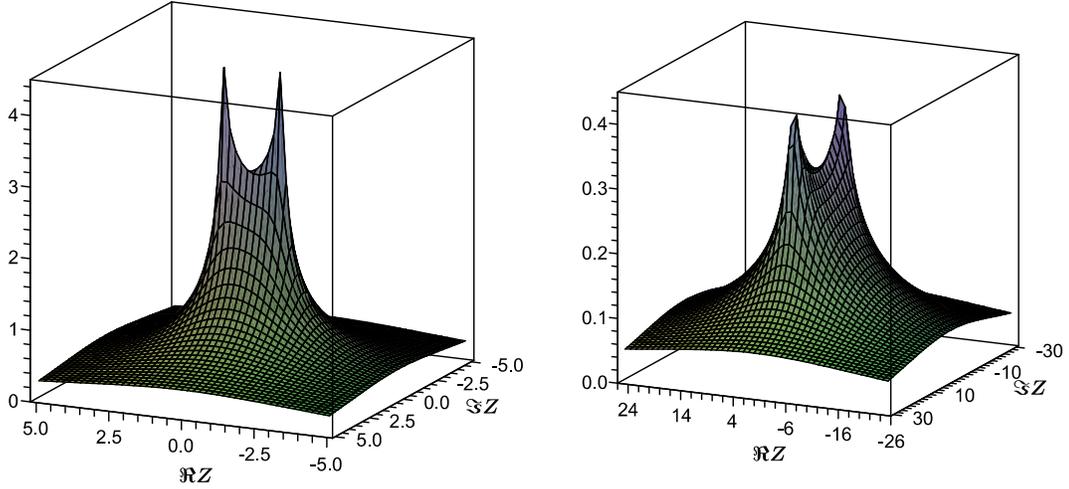


Figure 1. Plot of the passage time τ_p versus $\Re Z$ and $\Im Z$. Left panel: $\Re\Omega = 1$, $\Im\Omega = 0.1$. Right panel: $\Re\Omega = 1$, $\Im\Omega = 10$. As can be seen from the plots, the singularity occurs at the points $Z = \pm\Omega$ and $\tau_p = \pi/|\Omega|$ at the saddle point $Z = 0$.

3 Quantum evolution in the vicinity of the exceptional point

In the presence of exceptional points characterized by coalescence of eigenvalues and corresponding eigenvectors, non-Hermitian physics differs essentially from Hermitian physics [20, 25, 26]. Since for the Hermitian operators, the coalescence of eigenvalues results in different eigenvectors, exceptional points do not exist for Hermitian systems. The related degeneracies, referred to as ‘conical intersections’, are known also as ‘diaboloic points’. In the context of Berry phase, the diaboloic point is associated with the ‘fictitious magnetic monopole’ located at the diaboloic point [27, 28]. In turn, the exceptional points are associated with the ‘fictitious complex magnetic monopoles’ [29].

For a two-dimensional system (2.4), the eigenvalues coalesce when $\Omega = 0$. Writing the complex vector Ω as $\Omega = \mathbf{r} - i\delta$, where \mathbf{r} and δ are real, one can recast the effective Hamiltonian of equation (2.4) in the form

$$H_{ef} = \frac{\lambda_0}{2}\mathbf{1} + \frac{1}{2}\mathbf{r} \cdot \boldsymbol{\sigma} - \frac{i}{2}\delta \cdot \boldsymbol{\sigma}. \quad (3.1)$$

As can be seen, the degeneracy points are defined by the following equation:

$$r^2 - \delta^2 - 2ir\delta \cos \gamma = 0, \quad (3.2)$$

where γ denotes the angle between δ and $\mathbf{r} = (x, y, z)$. Furthermore, it is convenient to use cylindrical coordinates (ρ, z, φ) , in which $\Omega = \sqrt{\rho^2 + z^2 - \delta^2 - 2iz\delta}$, and equation (3.2) reads: $\rho^2 + z^2 - \delta^2 - 2iz\delta = 0$. There are two solutions to this equation. The first one is $\rho = z = \delta = 0$, and the corresponding degeneracy point located at the origin of coordinates is the *diaboloic point*. The second solution, given by $z = 0$ and $\rho^2 - \delta^2 = 0$, defines the *exceptional point*. Generally speaking, the exceptional ‘point’, being realized as a cone in the four-parameter space (x, y, z, δ) , is not a point in the convenient sense, and all possible one(zero)-dimensional cases can be obtained by the conic sections. For instance, for a fixed δ , we obtain the exceptional point as a circle of radius $|\delta|$, lying in the plane $z = 0$. It should be noted that in the recent literature, the term ‘exceptional point’ is applied not only to the particular case when the exceptional ‘point’ is indeed a point but to the general case when the non-Hermitian degeneracy is realized as a submanifold in the parameter space [30, 31, 32, 33]. In the following, we will stick to this more general interpretation of the exceptional point.

As follows from equations (2.9)–(2.13), at the exceptional point, the quantum evolution is described by

$$|\psi(t)\rangle = \left(1 - \frac{\delta t}{2(1 - \cos \frac{\alpha}{2})}\right) e^{-i\lambda_0 t} |\psi_i\rangle + \frac{\delta t}{2(1 - \cos \frac{\alpha}{2})} e^{-i\lambda_0 t} |\psi_f\rangle.$$

Without loss of generality, we further assume $\delta > 0$. The computation of the time evolution at the exceptional point then yields

$$\tau = \frac{2}{\delta} \left| \left(1 - \cos \frac{\alpha}{2}\right) \right|.$$

In what follows, imposing the eigenvalue constraint as $|\Omega| = \text{const}$, we restrict ourselves by consideration of the orthogonal initial $|\psi_i\rangle$ and final states $|\psi_f\rangle$. Substituting $\alpha = \pi$ into equation (2.13) and taking into account equations (2.15), (2.16), we obtain

$$|\psi(t)\rangle = \left(\cos \frac{\Omega t}{2} - i \cos \theta \sin \frac{\Omega t}{2} \right) e^{-i\lambda_0 t} |\psi_i\rangle + \sin \theta \sin \frac{\Omega t}{2} e^{-i\lambda_0 t} |\psi_f\rangle. \quad (3.3)$$

Now recalling that $\cos \theta = (z - i\delta)/\Omega$ and setting $z = 0$, we observe that in the vicinity of the exceptional point, there are two different regimes, depending on the relation between ρ and δ .

For $\rho > \delta$, we have $\Omega = \sqrt{\rho^2 - \delta^2} > 0$, and inserting $\cos \theta = (z - i\delta)/\Omega$ into equation (3.3), we obtain

$$|\psi(t)\rangle = \left(\cos \frac{\Omega_0 t}{2} - \frac{\delta}{\Omega_0} \sin \frac{\Omega_0 t}{2} \right) e^{-i\lambda_0 t} |\psi_i\rangle + \frac{\rho}{\Omega_0} \sin \frac{\Omega_0 t}{2} e^{-i\lambda_0 t} |\psi_f\rangle,$$

where $\Omega_0 = |\rho^2 - \delta^2|^{1/2}$ denotes the Rabi frequency. If, in addition, $\Im \lambda_0 = 0$, then the eigenvalues of the non-Hermitian Hamiltonian (3.1) are real, and we obtain the \mathcal{PT} -symmetric Hamiltonian widely discussed in the recent literature in connection with the quantum brachistochrone problem [2, 5, 7, 8, 9, 10, 11, 12, 34, 35]. Then, applying (2.18), we get (see Fig. 2)

$$\tau_p = \frac{2}{\Omega_0} \arctan \left(\frac{\Omega_0}{\delta} \right). \quad (3.4)$$

As can be shown, the passage time τ_p is bounded above and below as follows: $2/\sqrt{\Omega_0^2 + \delta^2} < \tau_p < \min\{\pi/\Omega_0, 2/\delta\}$. This improves the estimation of the passage time obtained in [2].

As follows from equation (3.4), for $\Omega_0 \gg \delta$, the passage time has the same behavior as for the equivalent Hermitian system: $\tau_p \sim \pi/\Omega_0$, and if $\Omega_0 \ll \delta$, one obtains $\tau_p \sim 2/\delta$. At the exceptional point $\Omega_0 = 0$ (and $\rho = \delta$):

$$|\psi(t)\rangle = \left(1 - \frac{\delta t}{2}\right) e^{-i\lambda_0 t} |\psi_i\rangle + \frac{\delta t}{2} e^{-i\lambda_0 t} |\psi_f\rangle.$$

From here, we get $\tau_p = 2/\delta$ and $\tau_p \rightarrow 0$ if $\delta \rightarrow \infty$. Thus, in contrast to the Hermitian systems, the passage time may be arbitrarily small; this, of course, implies that ρ, δ becomes large. This is possible due to the hyperbolic nature of Ω_0 considered as function of ρ , and δ and agrees with the results obtained in Bender et al. [2]. As a result, the equivalent Hermitian system in the limit $\Omega_0 \rightarrow 0$ yields $\tau_p = \pi/\Omega_0 \rightarrow \infty$.

Now assuming $\rho < \delta$, we have $\Omega = i\Omega_0$ and

$$|\psi(t)\rangle = \left(\cosh \frac{\Omega_0 t}{2} - \frac{\delta}{\Omega_0} \sinh \frac{\Omega_0 t}{2} \right) e^{-i\lambda_0 t} |\psi_i\rangle + \frac{\rho}{\Omega_0} \sinh \frac{\Omega_0 t}{2} e^{-i\lambda_0 t} |\psi_f\rangle.$$

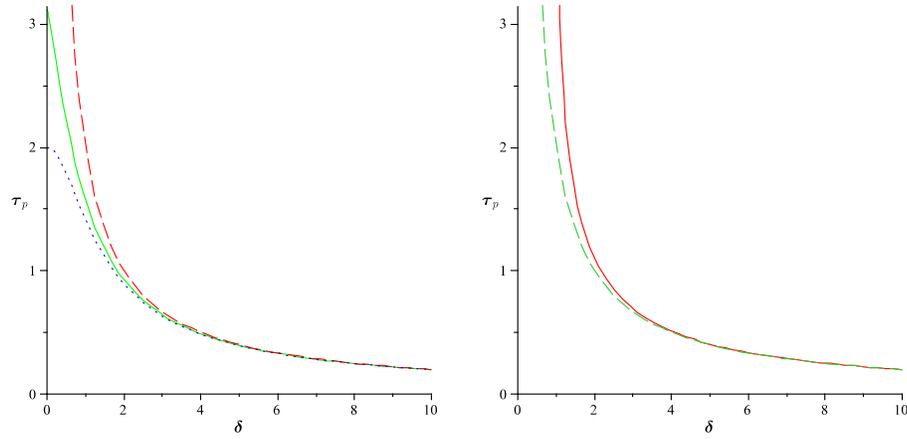


Figure 2. Plot of passage time τ_p vs. $x = \delta$ ($\Omega_0 = 1$). Left panel: $\tau_p = (2/\Omega_0) \arctan(\Omega_0/\delta)$ (solid green line), $\tau = 2/\sqrt{\Omega_0^2 + \delta^2}$ (dotted blue line) and $\tau = 2/\delta$ (dashed red line). Right panel: $\tau_p = (2/\Omega_0) \tanh^{-1}(\Omega_0/\delta)$ (solid red line) and $\tau = 2/\delta$ (dashed green line). It is seen that in both cases, τ_p is asymptotically approximated by $\tau = 2/\delta$.

Computation of the passage time yields

$$\tau_p = \frac{2}{\Omega_0} \tanh^{-1} \left(\frac{\Omega_0}{\delta} \right). \quad (3.5)$$

As seen in Fig. 2, the passage time is bounded below as follows: $\tau_p > 2/\delta$, and it follows from equation (3.5) that $\tau_p \sim 2/\delta$ in the limit $\Omega_0 \ll \delta$. At the exceptional point, we obtain the same result as above: $\tau_p = 2/\delta$ and

$$|\psi(t)\rangle = \left(1 - \frac{\delta t}{2} \right) e^{-i\lambda_0 t} |\psi_i\rangle + \frac{\delta t}{2} e^{-i\lambda_0 t} |\psi_f\rangle.$$

As an illustrative example, we consider a two-level dissipative system driven by a periodic electromagnetic field $\mathbf{E}(t) = \Re(\mathcal{E}(t) \exp(i\nu t))$. In the rotating wave approximation, after removing the explicit time dependence of the Hamiltonian and the average effect of the decay terms, the Schrödinger equation reads [36, 37]

$$i \begin{pmatrix} \dot{u}_1 \\ \dot{u}_2 \end{pmatrix} = \frac{1}{2} \begin{pmatrix} -i\lambda + \Delta - i\delta & 2V^* \\ 2V & -i\lambda - \Delta + i\delta \end{pmatrix} \begin{pmatrix} u_1 \\ u_2 \end{pmatrix}, \quad (3.6)$$

where $\lambda = (\gamma_a + \gamma_b)/2$, with γ_a, γ_b being decay rates for upper and lower levels, respectively, $\Delta = E_a - E_b - \nu$, $\delta = (\gamma_a - \gamma_b)/2$, $V = \boldsymbol{\mu} \cdot \boldsymbol{\mathcal{E}}$, and $\boldsymbol{\mu}$ is the dipole matrix element.

The choice $\boldsymbol{\mathcal{E}}(t) = \boldsymbol{\mathcal{E}}_0 \exp(i\omega t)$ yields $V(t) = V_0 \exp(i\omega t)$, where $V_0 = \boldsymbol{\mu} \cdot \boldsymbol{\mathcal{E}}_0$, and we assume further that $V_0 > 0$. The solution of equation (3.6) with this choice of $\boldsymbol{\mathcal{E}}$ can be written as

$$|u(t)\rangle = C_1(t) e^{-i(\omega - i\lambda)t/2} |u_\uparrow\rangle + C_2(t) e^{i(\omega + i\lambda)t/2} |u_\downarrow\rangle,$$

where $|u_\uparrow\rangle = \begin{pmatrix} 1 \\ 0 \end{pmatrix}$ and $|u_\downarrow\rangle = \begin{pmatrix} 0 \\ 1 \end{pmatrix}$, denote the up/down states, respectively,

$$|C(t)\rangle = \begin{pmatrix} C_1(t) \\ C_2(t) \end{pmatrix} = \begin{pmatrix} \cos \frac{\Omega t}{2} - i \cos \theta \sin \frac{\Omega t}{2} & -i \sin \theta \sin \frac{\Omega t}{2} \\ -i \sin \theta \sin \frac{\Omega t}{2} & \cos \frac{\Omega t}{2} + i \cos \theta \sin \frac{\Omega t}{2} \end{pmatrix} \begin{pmatrix} C_1(0) \\ C_2(0) \end{pmatrix},$$

and we set $\cos \theta = (\Delta - \omega - i\delta)/\Omega$, $\Omega = (\rho^2 + (\Delta - \omega - i\delta)^2)^{1/2}$, $\rho = 2V_0$.

As can be shown, the state $|C(t)\rangle$ satisfies the Schrödinger equation

$$i\frac{\partial|C\rangle}{\partial t} = H_r|C\rangle,$$

written in a co-rotating reference frame, where the Hamiltonian H_r takes the form

$$H_r = \frac{\Omega}{2} \begin{pmatrix} \cos\theta & \sin\theta \\ \sin\theta & -\cos\theta \end{pmatrix}.$$

Let $|u(t)\rangle$ be the solution of equation (3.6) with the initial state at $t = 0$ taken as $|u_\uparrow\rangle$ and the final state of the system at a later time t being $|u_\uparrow\rangle$ or $|u_\downarrow\rangle$. Then following Baker [38], we compute the probability $P_{\uparrow\uparrow}$ ($P_{\downarrow\uparrow}$) that the system is in the state $|u_\uparrow\rangle$ ($|u_\downarrow\rangle$), respectively:

$$P_{\uparrow\uparrow} = \left| \cos\frac{\Omega t}{2} - i\cos\theta\sin\frac{\Omega t}{2} \right|^2 e^{-\lambda t}, \quad P_{\downarrow\uparrow} = \left| \sin\theta\sin\frac{\Omega t}{2} \right|^2 e^{-\lambda t}.$$

For the resonance frequencies, $\omega = \Delta$, we have $\Omega = (\rho^2 - \delta^2)^{1/2}$, and $\cos\theta = -i\delta/\Omega_0$. There are two different regimes, depending on the relation between ρ and δ . For $\rho > \delta$, we have a *coherent* tunneling process

$$P_{\uparrow\uparrow} = e^{-\lambda t} \left(\cos\frac{\Omega_0 t}{2} - \frac{\delta}{\Omega_0} \sin\frac{\Omega_0 t}{2} \right)^2, \quad P_{\downarrow\uparrow} = e^{-\lambda t} \frac{\rho^2}{\Omega_0^2} \sin^2\frac{\Omega_0 t}{2}, \quad (3.7)$$

where $\Omega_0 = |\rho^2 - \delta^2|^{1/2}$ denotes the Rabi frequency. On the other hand, for $\rho < \delta$, there is *incoherent* tunneling

$$P_{\uparrow\uparrow} = e^{-\lambda t} \left(\cosh\frac{\Omega_0 t}{2} - \frac{\delta}{\Omega_0} \sinh\frac{\Omega_0 t}{2} \right)^2, \quad P_{\downarrow\uparrow} = e^{-\lambda t} \frac{\rho^2}{\Omega_0^2} \sinh^2\frac{\Omega_0 t}{2}. \quad (3.8)$$

At the exceptional point, $\Omega_0 = 0$, and both regimes yield

$$P_{\uparrow\uparrow} = \left(1 - \frac{\delta t}{2} \right)^2 e^{-\lambda t}, \quad P_{\downarrow\uparrow} = \left(\frac{\delta t}{2} \right)^2 e^{-\lambda t}.$$

This is in accordance with the results obtained by Stafford and Barrett and Dietz et al. [39, 40]. Moreover, the decay behavior predicted by equations (3.7)–(3.8) has been observed in the experiment with a dissipative microwave billiard [40].

We find that for $\rho > \delta$, the passage time required to transform the state $|u_\uparrow\rangle$ into the state $|u_\downarrow\rangle$ is given by $\tau_p = (2/\omega_0) \arctan(\omega_0/\delta)$. Similarly, for $\rho < \delta$, we obtain $\tau_p = (2/\omega_0) \tanh^{-1}(\omega_0/\delta)$. At the exceptional point, for both regimes, the passage time is found to be $\tau_p = 2/\delta$ (see Fig. 2).

4 Fubini–Study metric on the complex Bloch sphere and the brachistochrone problem

In this section, we will consider the quantum non-Hermitian brachistochrone problem in the context of the geometric approach developed by Anandan and Aharonov [41]. Let $|\psi(t)\rangle$ and $\langle\tilde{\psi}(t)|$ satisfy the Schrödinger equation and its adjoint equation, respectively:

$$i\frac{\partial}{\partial t}|\psi(t)\rangle = H|\psi(t)\rangle, \quad -i\frac{\partial}{\partial t}\langle\tilde{\psi}(t)| = \langle\tilde{\psi}(t)|H,$$

where H is a non-Hermitian Hamiltonian, and the standard normalization is held, $\langle \tilde{\psi}(t) | \psi(t) \rangle = 1$. We define the complex energy variance as

$$\Delta E^2 = \langle \tilde{\psi} | H^2 | \psi \rangle - (\langle \tilde{\psi} | H | \psi \rangle)^2.$$

Now, applying the Taylor expansion to $|\psi(t + dt)\rangle$ and using equation (4) and its time derivative, we obtain

$$\langle \tilde{\psi}(t) | \psi(t + dt) \rangle^2 = 1 - \Delta E^2 dt^2 + \mathcal{O}(dt^3).$$

Then, introducing the complex metric as $ds^2 = 4(1 - \langle \tilde{\psi}(t) | \psi(t + dt) \rangle^2)$, we obtain

$$ds^2 = 4\Delta E^2 dt^2 = 4ds_{\text{FS}}^2, \quad (4.1)$$

where

$$ds_{\text{FS}}^2 = \langle d\tilde{\psi} | (1 - P) | d\psi \rangle \quad (4.2)$$

is a natural generalization of the Fubini–Study line element to the non-Hermitian quantum mechanics, $P = |\psi\rangle\langle\tilde{\psi}|$ being the projection operator to the state $|\psi\rangle$. It is easy to show that the complex-valued metric (4.2) is gauge invariant, i.e., it does not depend on the particular choice of the complex phase defined by the map: $|\psi\rangle \rightarrow e^{i\alpha}|\psi\rangle$, $\langle\tilde{\psi}| \rightarrow e^{-i\alpha}\langle\tilde{\psi}|$, $\alpha \in \mathbb{C}$. We define the distance between two given states $|\psi_0\rangle$ and $|\psi_1\rangle$ as

$$s = 2 \int_{\mathcal{C}} |\Delta E(t)| dt, \quad (4.3)$$

where the integration is performed along a given curve \mathcal{C} connecting $|\psi_0\rangle$ and $|\psi_1\rangle$.

In the two-dimensional case, the complex-valued Fubini–Study element has a nice geometric interpretation. We define a complex distance between two nearby Bloch vectors $\mathbf{n}(x)$ and $\mathbf{n}(x + dx)$ by

$$\Delta^2(x, x + dx) = 1 - \mathbf{n}(x) \cdot \mathbf{n}(x + dx).$$

Then, applying Taylor expansion,

$$\mathbf{n}(x + dx) = \mathbf{n}(x) + \partial_i \mathbf{n}(x) dx^i + \frac{1}{2} \partial_i \partial_j \mathbf{n}(x) dx^i dx^j + \dots,$$

and using $\mathbf{n} \cdot \mathbf{n} = 1$, we obtain, up to second-order terms,

$$\mathbf{n}(x + dx) = 1 - d\mathbf{n}(x) \cdot d\mathbf{n}(x).$$

This yields

$$ds^2 = d\mathbf{n} \cdot d\mathbf{n} = g_{ij} dx^i dx^j, \quad (4.4)$$

where $g_{ij} = \partial_i \mathbf{n} \cdot \partial_j \mathbf{n}$, and the length of any curve \mathcal{C} on S_c^2 is given by

$$L(\mathcal{C}) = \int_{\mathcal{C}} |\sqrt{d\mathbf{n} \cdot d\mathbf{n}}|.$$

Denoting $\mathbf{n} = (\sin \zeta \cos \nu, \sin \zeta \sin \nu, \cos \zeta)$, where $\zeta, \nu \in \mathbb{C}$, we recast (4.4) as

$$ds^2 = d\zeta^2 + \sin^2 \zeta d\nu^2. \quad (4.5)$$

Finally, using the definition of the complex Bloch vector $\mathbf{n} = \langle \tilde{\psi} | \boldsymbol{\sigma} | \psi \rangle$, we find that the metric on the complex Bloch sphere S_c^2 can be written as

$$ds^2 = d\mathbf{n} \cdot d\mathbf{n} = 4ds_{\text{FS}}^2 = 4\langle d\tilde{\psi} | (1 - P) | d\psi \rangle,$$

where $ds_{\text{FS}}^2 = \langle d\tilde{\psi} | (1 - P) | d\psi \rangle$ is the Fubini–Study line element.

Strictly speaking, g_{ij} being a complex-valued tensor does not define a proper metric on the complex Bloch sphere. However, the advantage of this definition is that, contrary to the Kähler metric, the complex “metric” (4.4) is invariant under the gauge transformation $|\psi\rangle \rightarrow e^{i\alpha}|\psi\rangle$, $\langle \tilde{\psi} | \rightarrow e^{-i\alpha}\langle \tilde{\psi} |$, where $\alpha \in \mathbb{C}$.

For the Hamiltonian (2.20), the straightforward computation yields

$$2\Delta E = \Omega \sin \theta, \tag{4.6}$$

and, using equations (2.18) and (4.3), we obtain

$$L = 2|\Delta E|\tau = \left| 2\sin \theta \arctan \left(\frac{i \sin^2 \frac{\alpha}{2}}{\cos \frac{\alpha}{2} \sqrt{\cos^2 \theta - \sin^2 \frac{\alpha}{2} - \cos \theta}} \right) \right|. \tag{4.7}$$

To compare our results with the \mathcal{PT} -symmetric model widely discussed in the literature (see, e.g., [6, 7, 8, 9, 10, 11, 12, 13, 14]), we choose $\theta = \pi/2 + i\eta$ and assume $\Im \lambda_0 = \Im \varphi = 0$ to write the effective Hamiltonian of equation (2.20) as

$$H_{\text{eff}} = \frac{1}{2} \begin{pmatrix} \lambda_0 & 0 \\ 0 & \lambda_0 \end{pmatrix} + \frac{\Omega}{2} \begin{pmatrix} -i \sinh \eta & e^{-i\varphi} \cosh \eta \\ e^{i\varphi} \cosh \eta & i \sinh \eta \end{pmatrix}. \tag{4.8}$$

Next, we set $\lambda_0 = r \cos \gamma$, $\Omega \sinh \eta = r \sin \gamma$, and $\Omega \cosh \eta = \omega$. Denoting $\delta = r \sin \gamma$, we obtain the Hamiltonian of the system in its conventional form (see, e.g., [2, 5]):

$$H_{\text{eff}} = \frac{1}{2} \begin{pmatrix} r e^{-i\gamma} & \omega e^{-i\varphi} \\ \omega e^{i\varphi} & r e^{i\gamma} \end{pmatrix},$$

where $\Omega = \sqrt{\omega^2 - r^2 \sin^2 \gamma} = \sqrt{\omega^2 - \delta^2}$. In this particular case, the complex Bloch sphere becomes the one-sheeted two-dimensional hyperboloid \mathbb{H}^2 with the indefinite metric given by [24]

$$ds^2 = \cosh^2 \rho d\nu^2 - d\rho^2, \tag{4.9}$$

where $-\infty < \rho < \infty$ and $0 \leq \nu < 2\pi$ are the inner parameters on \mathbb{H}^2 . It should be noted that the interval (4.9) can be obtained from equation (4.5) by the substitution $\zeta \rightarrow \pi/2 + i\rho$, and we assume $\Im \nu = 0$.

The amount of time τ required to evolve the initial state $|\psi_i\rangle$ into the final state $|\psi_f\rangle$ can be found from equation (2.18) by substituting $\theta = \pi/2 + i\eta$. The computation yields

$$\tau = \frac{2}{\Omega} \arctan \left(\frac{\sin^2 \frac{\alpha}{2}}{\cos \frac{\alpha}{2} \sqrt{\sinh^2 \eta + \sin^2 \frac{\alpha}{2} + \sinh \eta}} \right), \tag{4.10}$$

and from equation (4.7), we obtain

$$L = 2 \cosh \eta \arctan \left(\frac{\sin^2 \frac{\alpha}{2}}{\cos \frac{\alpha}{2} \sqrt{\sinh^2 \eta + \sin^2 \frac{\alpha}{2} + \sinh \eta}} \right).$$

To study the spin-flip, we choose the initial state as $|\psi_i\rangle = |u_\uparrow\rangle$. Then, substituting $\alpha = \pi$ into equation (4.10), we find the time interval

$$\tau_\downarrow = \frac{2}{\Omega} \arctan \frac{1}{\sinh \eta} = \frac{2}{\Omega} \arctan \frac{\Omega}{\delta}, \quad \tau_\uparrow = \frac{2\pi}{\Omega} - \frac{2}{\Omega} \arctan \frac{\Omega}{\delta},$$

necessary for the first spin flip from up to down and back. For all values $\Omega \in [0, \infty)$, we have $2/\delta \leq \tau_\downarrow \leq \pi/\Omega$. Thus, the passage time $\tau_p = \tau_\downarrow$ lies below the Anandan–Aharonov lower bound, $\tau_\downarrow \leq \pi/\Omega$ for a spin-flip evolution in a Hermitian system [42], and τ_p reaches its minimum value $\tau_{\min} = 2/\delta$ at the exceptional point. In addition, the total time for a spin-flip evolution, $|\uparrow\rangle \rightarrow |\downarrow\rangle \rightarrow |\uparrow\rangle$, remains invariant: $\tau = \tau_\downarrow + \tau_\uparrow = 2\pi/\Omega$. This is in accordance with the results of previous studies [34, 10].

Introducing the variable $\kappa = \arctan(\delta/\Omega) = \arctan(\sinh \eta)$, we reproduce our results in the more familiar form, widely known in the literature (see, e.g., [2, 3, 4, 10]):

$$\tau_\downarrow = \frac{\pi - 2\kappa}{\Omega}, \quad \tau_\uparrow = \frac{\pi + 2\kappa}{\Omega}.$$

We see that in the Hermitian limit, $\eta \rightarrow 0$ ($\delta \ll \Omega$) that implies $\kappa \rightarrow 0$, the passage time is given by $\tau_p = \pi/\Omega$. In the other limiting case $\eta \rightarrow \infty$ ($\Omega \ll \delta$), the angle κ approaches $\pi/2$, and τ_p tends to zero. In terms of variable κ , the relation $\Omega \cosh \eta = \omega$ becomes $\Omega = \omega \cos \kappa$. Furthermore, if the energy constraint $E_+ - E_- = \Omega$ is held fixed, in order to have the passage to the limit $\kappa \rightarrow \pi/2$, one must require $\omega \gg \Omega$. It then follows from the relation $\Omega^2 = \omega^2 - \delta^2 = \text{const}$ that one must require $\delta \gg \Omega$.

Similar consideration of the distance between the initial and the final states yields

$$L_p = 2 \cosh \eta \arctan \left(\frac{1}{\sinh \eta} \right) = \frac{\pi - 2\kappa}{\cos \kappa}. \quad (4.11)$$

In the limit $\kappa \rightarrow \pi/2$, we get $L_p \rightarrow 2$, and in the Hermitian limit, $\kappa \rightarrow 0$, we have $L_p \rightarrow \pi$. It then follows from equation (4.11), that the distance between $|u_\uparrow\rangle$ and $|u_\downarrow\rangle$, being measured on the one-sheeted two-dimensional hyperboloid \mathbb{H}^2 with the indefinite metric of equation (4.9), is bounded as follows: $2 \leq L_p \leq \pi$.

Inserting $\theta = \pi/2 + i\eta$ into (4.6), we obtain $2\Delta E = \Omega \cos \eta$. Then substituting this result into equation (4.1), we find that for the non-Hermitian Hamiltonian (4.8), the evolution speed $v = ds/dt$ is given by $v = \omega = \Omega \cosh \eta$. Hence, $v \geq v_g$, where $v_g = \Omega$ is the speed of the geodesic evolution [24]. Similar consideration of the quantum-mechanical system governed by the Hermitian Hamiltonian yields $v = \Omega \sin \theta$, and, obviously, $v \leq v_g$. This proves that non-Hermitian quantum mechanics can be faster than Hermitian quantum mechanics. Moreover, since for any complex angle θ , there exists θ_0 determined by the equation $\cos \Re \theta_0 = \sinh \Im \theta_0$ such that $v = |\Omega \sin \theta| \geq |\Omega| = v_g$ if $|\Im \theta| \geq |\Im \theta_0|$, this conclusion is applied to an arbitrary non-Hermitian Hamiltonian (see Fig. 3).

Our results are in agreement with those obtained previously by Bender et al. [2]. However, interpreting the obtained results for τ requires care. Indeed, as was pointed out by Mostafazadeh [7], to compare the time evolution for the non-Hermitian and Hermitian Hamiltonians and to conclude in which case the evolution is faster, one should not only impose the same set of constraints for both cases but also fix the geodesic distance between initial and final states.

In what follows, we consider in detail a spin-flip for the Hamiltonian (4.8) with two types of constraints: a) the energy variance $\Delta E^2 = \langle \tilde{\psi} | H^2 | \psi \rangle - (\langle \tilde{\psi} | H | \psi \rangle)^2$ held fixed [1], and b) the difference between the largest and smallest energy eigenvalues has a fixed value, $E_+ - E_- = \text{const}$ [2]. In terms of the parameters (Ω, ω) , the energy constraints are written as follows:

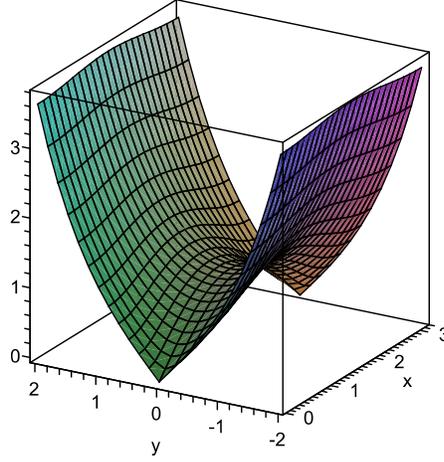


Figure 3. Plot of the evolution speed $|v|/|\Omega|$ versus $x = \Re\theta$ and $y = \Im\theta$. As can be observed, for any $\theta : 0 \leq \Re\theta \leq \pi$, there exists the angle θ_0 such that $|v|/|\Omega| \geq 1$, if $|\Im\theta| \geq |\Im\theta_0|$.

$2\Delta E = \omega$ and $E_+ - E_- = \Omega$. Then, using the relation $\cosh \eta = \omega/\Omega$, we find that the passage time can be recast as

$$\tau_p = \frac{2}{\Omega} \arctan \frac{\Omega}{\sqrt{\omega^2 - \Omega^2}}. \quad (4.12)$$

As can be observed, under the constraint $\omega = \text{const}$, the time evolution is bounded by $2/\omega \leq \tau_p \leq \pi/\omega$. Here, the passage time reaches its minimum value $\tau_{\min} = 2/\omega$ at the exceptional point $\Omega = 0$ ($\eta \rightarrow \infty$). The maximum, $\tau_{\max} = \pi/\omega$, is obtained for $\Omega = \omega$. Since $\Omega = \omega$ implies $\eta = 0$, this case corresponds to the Hermitian Hamiltonian. Next, we let the energy constraint be $E_+ - E_- = \Omega = \text{const}$. Then, as follows from equation (4.12), $\tau_p \leq \pi/\Omega$, and, just as above, the passage time achieves the maximum $\tau_{\max} = \pi/\Omega$ at the point $\omega = \Omega$ corresponding to the Hermitian Hamiltonian. For $\omega \gg \Omega$, we obtain $\tau_p \approx 2/\omega$, and τ_p vanishes in the limit $\omega \rightarrow \infty$.

Similar consideration of the distance between the initial and final states yields

$$L_p = \omega\tau = \frac{2\omega}{\Omega} \arctan \frac{\Omega}{\sqrt{\omega^2 - \Omega^2}}$$

and, under the constraint $\omega = \text{const}$, we have $2 \leq L \leq \pi$. The upper bound $L_{\max} = \pi$, being identical to the geodesic distance between the initial and final states defined either on the Bloch sphere S^2 or on the one-sheeted hyperboloid \mathbb{H}^2 [24], is thus achieved for the Hermitian Hamiltonian ($\Omega = \omega$). The lower bound, $L_{\min} = 2$, is obtained at the exceptional point ($\Omega = 0$). This agrees with our general conclusions regarding the behavior of the system in the vicinity of the exceptional point (see Section 3). Next, imposing the constraint $\Omega = \text{const}$, we obtain the same result $2 \leq L_p \leq \pi$. The minimum of L_p corresponds to the limit $\omega \gg \Omega$, and, just as above, the upper bound is achieved for the Hermitian Hamiltonian ($\omega = \Omega$).

We now turn our discussion to the recent controversy around the possibility of achieving faster evolution in a quantum-mechanical system governed by a non-Hermitian \mathcal{PT} -symmetric Hamiltonian as compared to the equivalent Hermitian system [2, 6, 7, 9]. It should be noted that the critique of the results obtained in [2] is essentially based on the following theorem [7]: *The lower bound on the travel time (upper bound on the speed) of unitary evolutions is a universal quantity, independent of whether the evolution is generated by a Hermitian or a non-Hermitian Hamiltonian.* Analyzing the proof of this theorem, one can see that it is founded on the following assumption: the minimal travel time is realized by quantum evolution along the geodesic path in the Hilbert space joining initial and final states.

This is true in the case of the Hermitian Hamiltonian; however, for a non-Hermitian Hamiltonian, the situation is quite different [24]. Let \mathbf{n}_i and \mathbf{n}_f denote antipodal states on the Bloch sphere S^2 , and \mathbf{m}_i and \mathbf{m}_f denote corresponding (antipodal) states on the one-sheeted hyperboloid \mathbb{H}^2 . Then, the geodesic distance L_g between \mathbf{n}_i and \mathbf{n}_f calculated over the Bloch sphere S^2 is identical to the geodesic distance between \mathbf{m}_i and \mathbf{m}_f computed over the one-sheeted hyperboloid \mathbb{H}^2 (for detailed calculations, see [24]). Moreover, under the same set of constraints, the amount of time τ_g required to evolve \mathbf{n}_i into \mathbf{n}_f on the Bloch sphere is equal to that required to evolve \mathbf{m}_i into \mathbf{m}_f on \mathbb{H}^2 by geodesic evolution. This is in accordance with the conclusions made by Mostafazadeh in [7]. However, as we have shown above, in the case of the Hermitian Hamiltonian, τ_g is the lower bound on the time evolution, and, for the non-Hermitian Hamiltonian, it yields only the upper bound on the time evolution. Thus, in non-Hermitian quantum mechanics, the evolution of a system is indeed faster than in Hermitian quantum mechanics, subject to the same energy constraint.

5 Conclusion

In this paper, we considered the non-Hermitian quantum brachistochrone problem for the generic non-Hermitian Hamiltonian and focused our attention on the geometric aspects of the problem. We demonstrated that for the generic non-Hermitian Hamiltonian, the quantum brachistochrone problem can be effectively formulated on the complex Bloch sphere S_c^2 , enabling the latter to be considered as the quantum phase space of the related two-level system. In particular, for the effective non-Hermitian Hamiltonian with a real energy spectrum, the corresponding quantum phase space is represented by the one-sheeted hyperboloid \mathbb{H}^2 .

As noted in [4, 5], in the Hermitian quantum brachistochrone problem, the lower bound on the time evolution interval τ , being proportional to the geodesic distance between the initial and final states on the Bloch sphere, is determined using the Fubini-Study metric on $\mathbb{C}P^1 \cong S^2$. We have shown that the geodesic distance L_g between antipodal states \mathbf{n}_i and \mathbf{n}_f calculated over the Bloch sphere S^2 is identical to the geodesic distance between corresponding antipodal states \mathbf{m}_i and \mathbf{m}_f calculated over the one-sheeted hyperboloid \mathbb{H}^2 . Moreover, the amount of time τ_g required to evolve \mathbf{n}_i into \mathbf{n}_f on the Bloch sphere is equal to that required to evolve \mathbf{m}_i into \mathbf{m}_f on \mathbb{H}^2 by the geodesic evolution. This is in accordance with the conclusions made in [7]. However, in the case of the Hermitian Hamiltonian, τ_g is the lower bound on the time evolution, and, for the non-Hermitian Hamiltonian, it yields only the upper bound on the time evolution. Furthermore, while for a quantum-mechanical system governed by the Hermitian Hamiltonian the evolution speed v is bounded by $v \leq v_g$, where v_g is the speed of the geodesic evolution, for a non-Hermitian quantum system with the same energy constraint, we have $v \geq v_g$. This proves that in non-Hermitian quantum mechanics, evolution can be faster than in Hermitian quantum mechanics [2].

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